

483rd WE-Heraeus Seminar  
**Extrasolar Planets:**  
**Towards Comparative Planetology beyond the Solar System**

**Tides  
on Earth,  
in the Solar System  
and Beyond.....**

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IDENTIFYING THE "PROGRESSIVE" HORSESHOE CRAB



# TIDE CLOCK HIGH LOW



# TIDE CLOCK

## HIGH LOW

**tide levels for Hamburg, Sankt Pauli:**

**Tuesday, 7th June 2011 (CEST)**

Low tide	04:03	0.3 m
High tide	09:07	3.9 m
Low tide	16:07	0.4 m
High tide	21:16	4.3 m

Bundesamt für Seeschifffahrt und Hydrographie



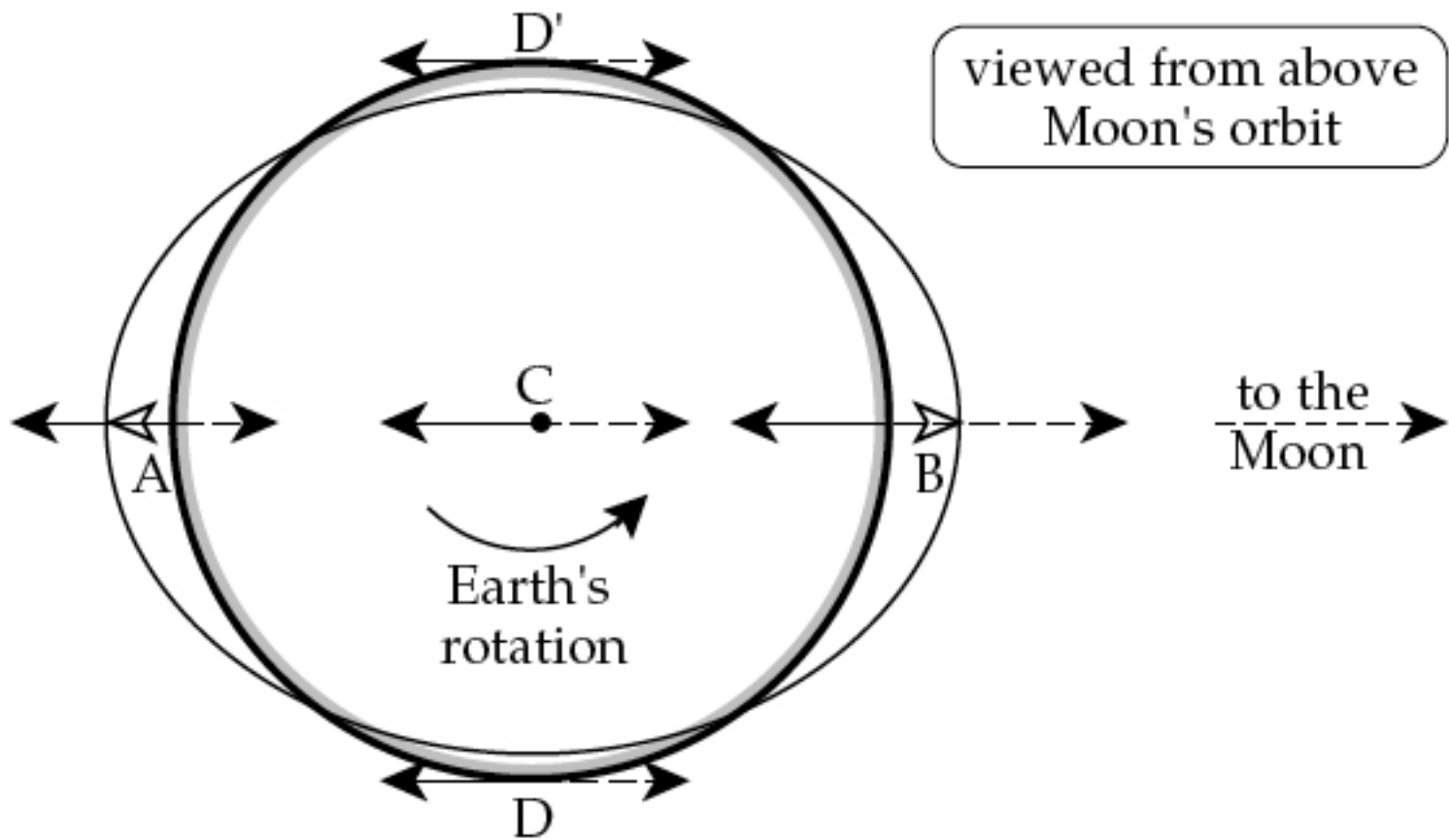
Mont St. Michel, Bretagne



Rosetta\_CI

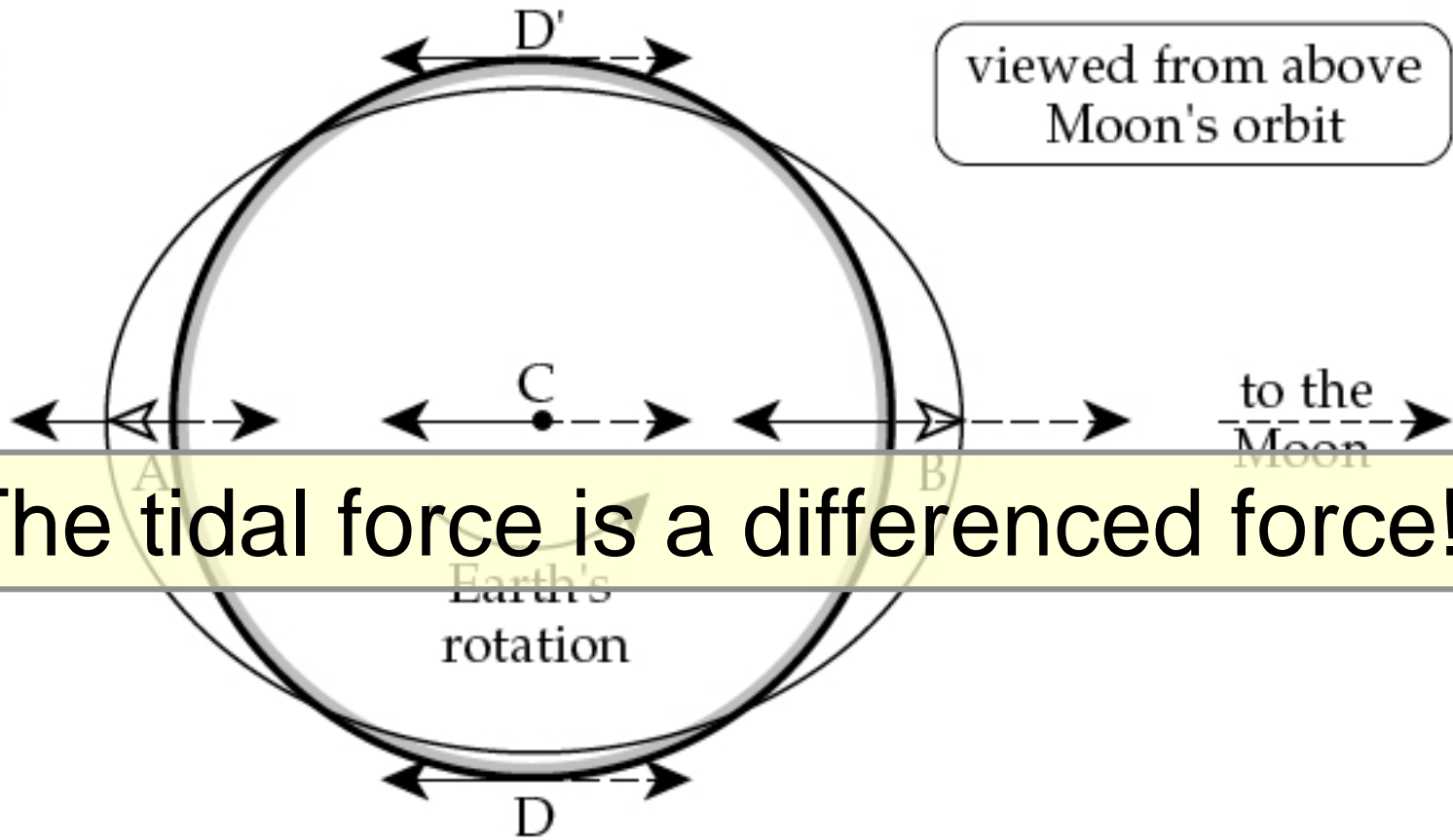


(a)



- $a_L$  ← constant centrifugal acceleration
- $a_G$  - - - → variable lunar gravitation
- $a_T$  → residual tidal acceleration

(a)

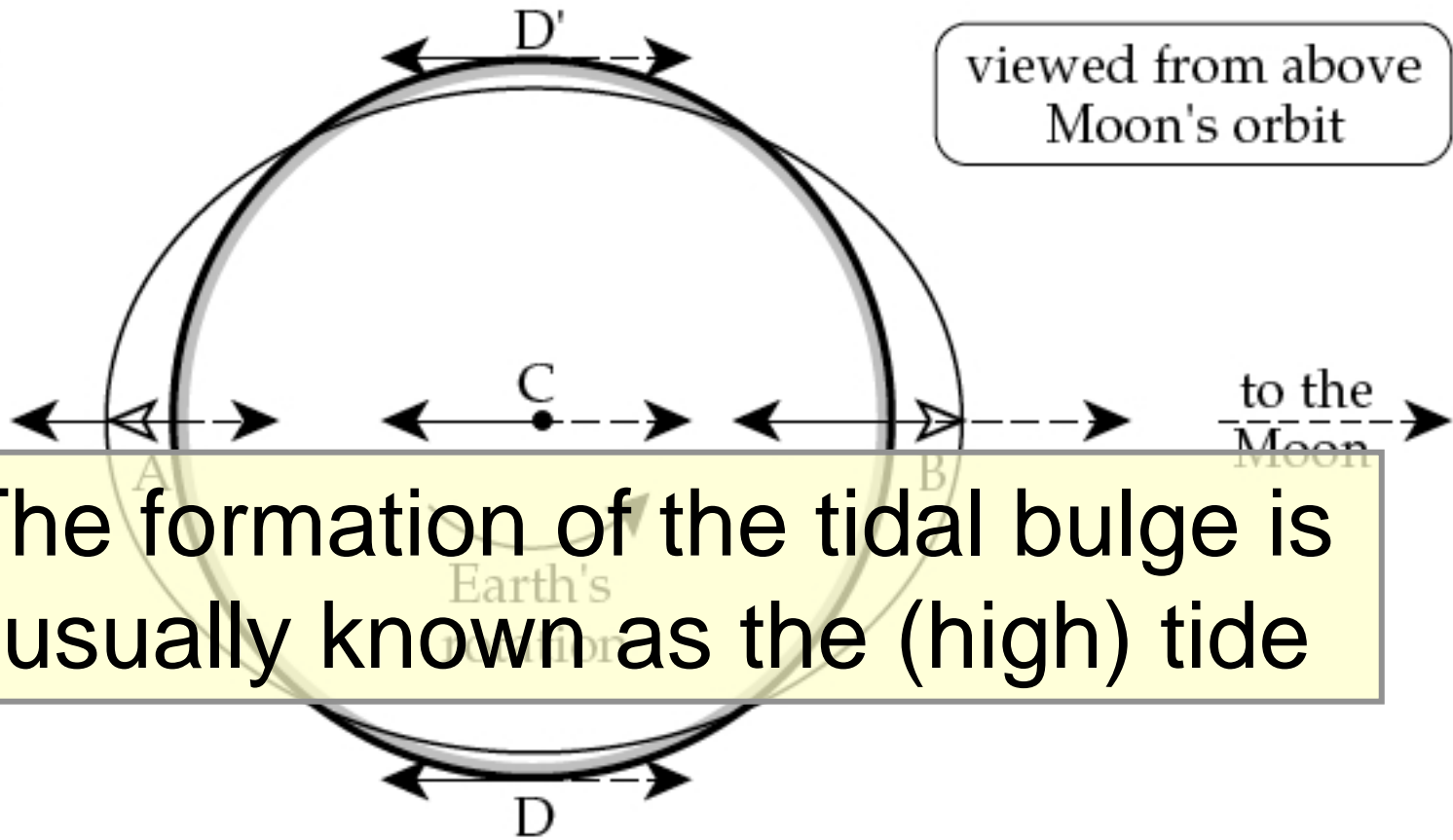


viewed from above  
Moon's orbit

The tidal force is a differenced force!

- $a_L$  ← constant centrifugal acceleration
- $a_G$  - - - → variable lunar gravitation
- $a_T$  → residual tidal acceleration

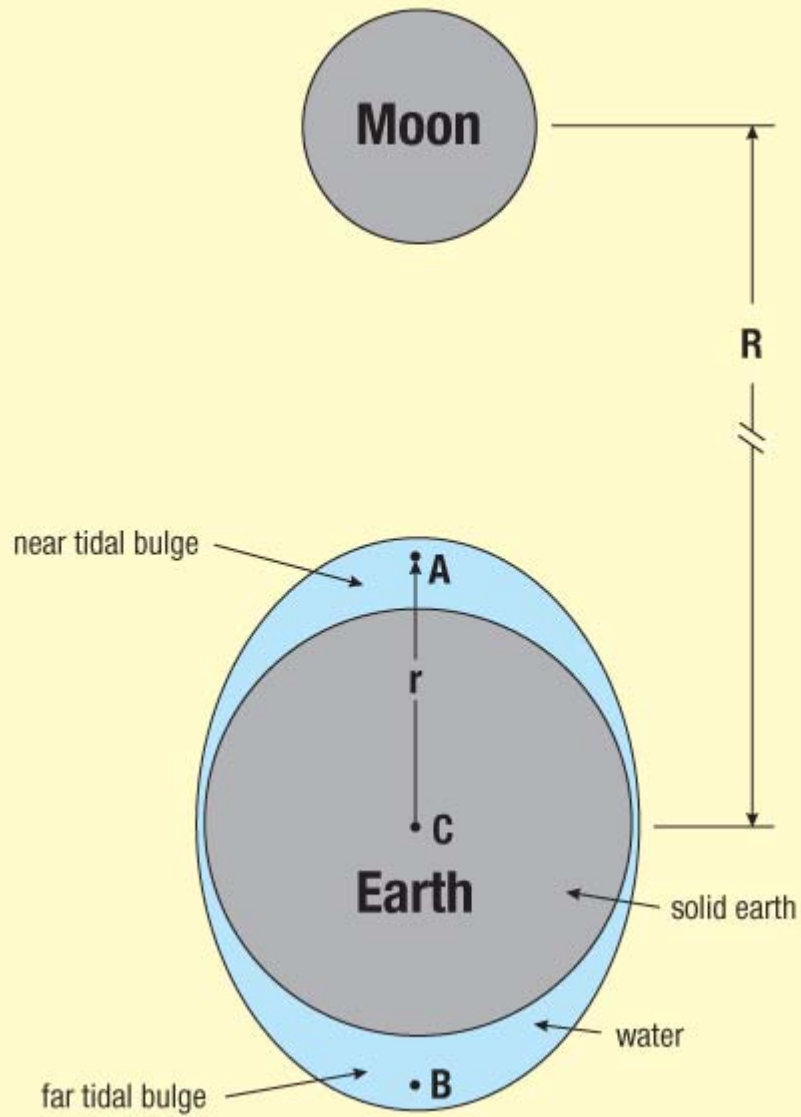
(a)



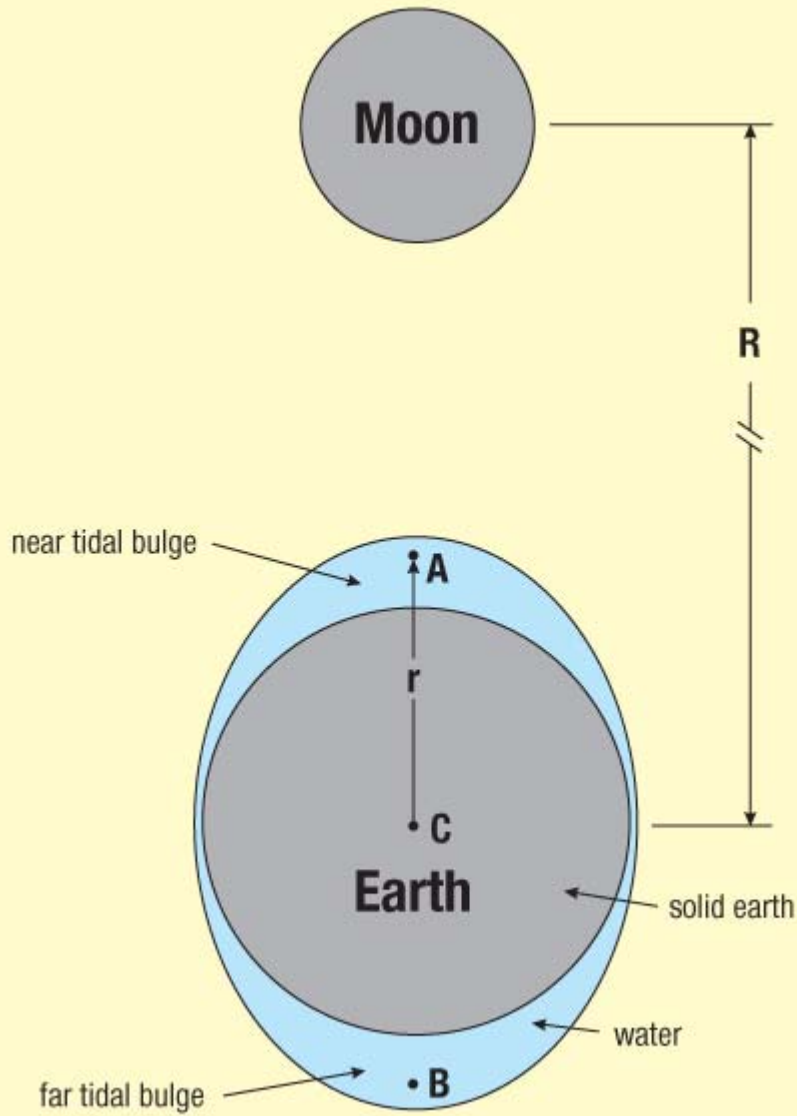
The formation of the tidal bulge is usually known as the (high) tide

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- $a_G$  - - - → variable lunar gravitation
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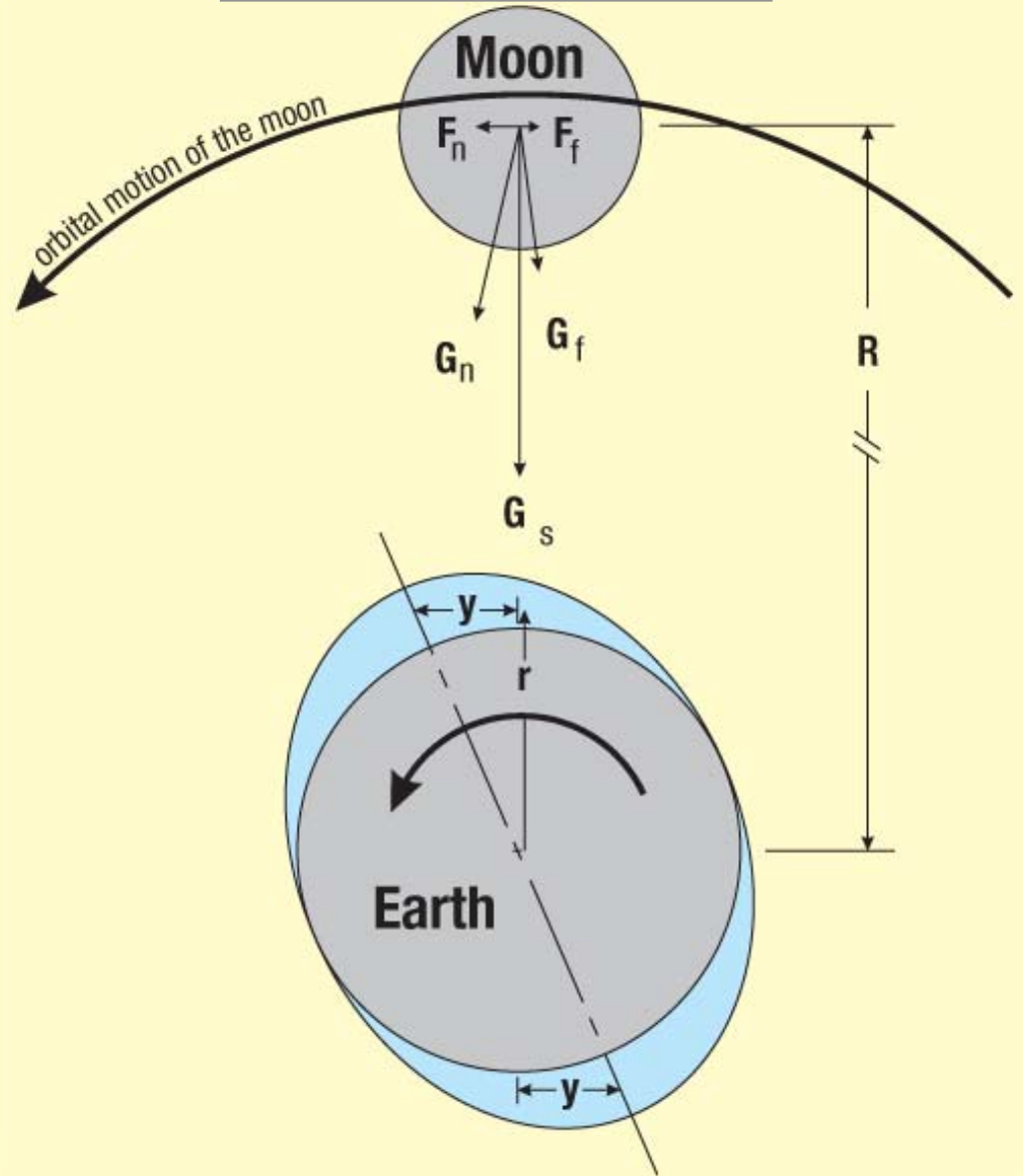
# Revolution without rotation



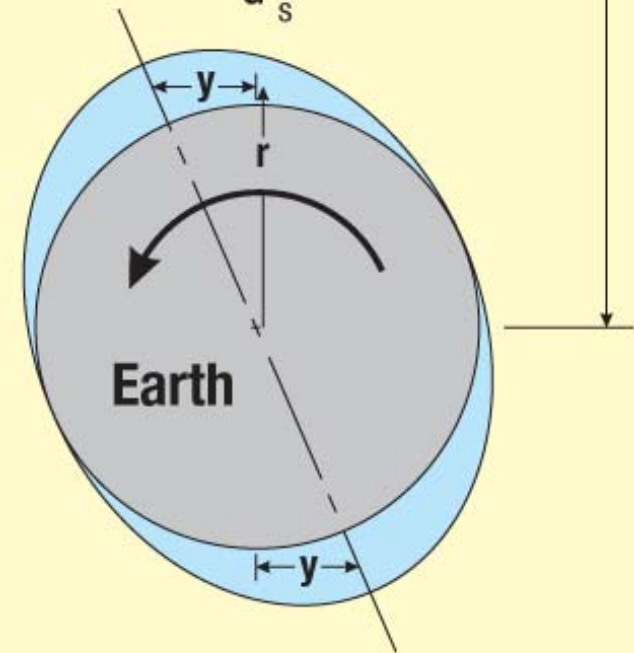
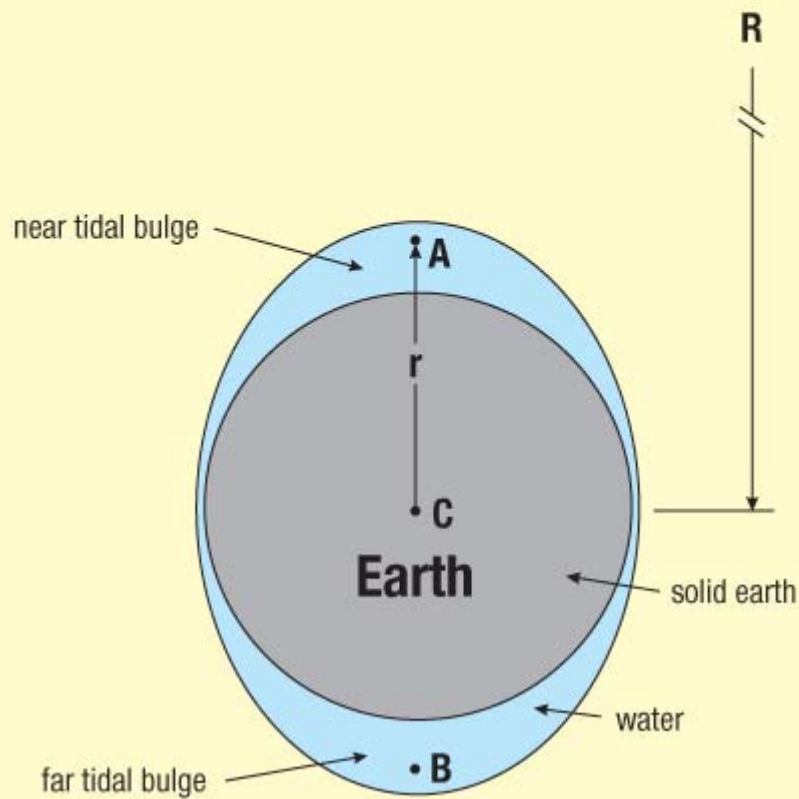
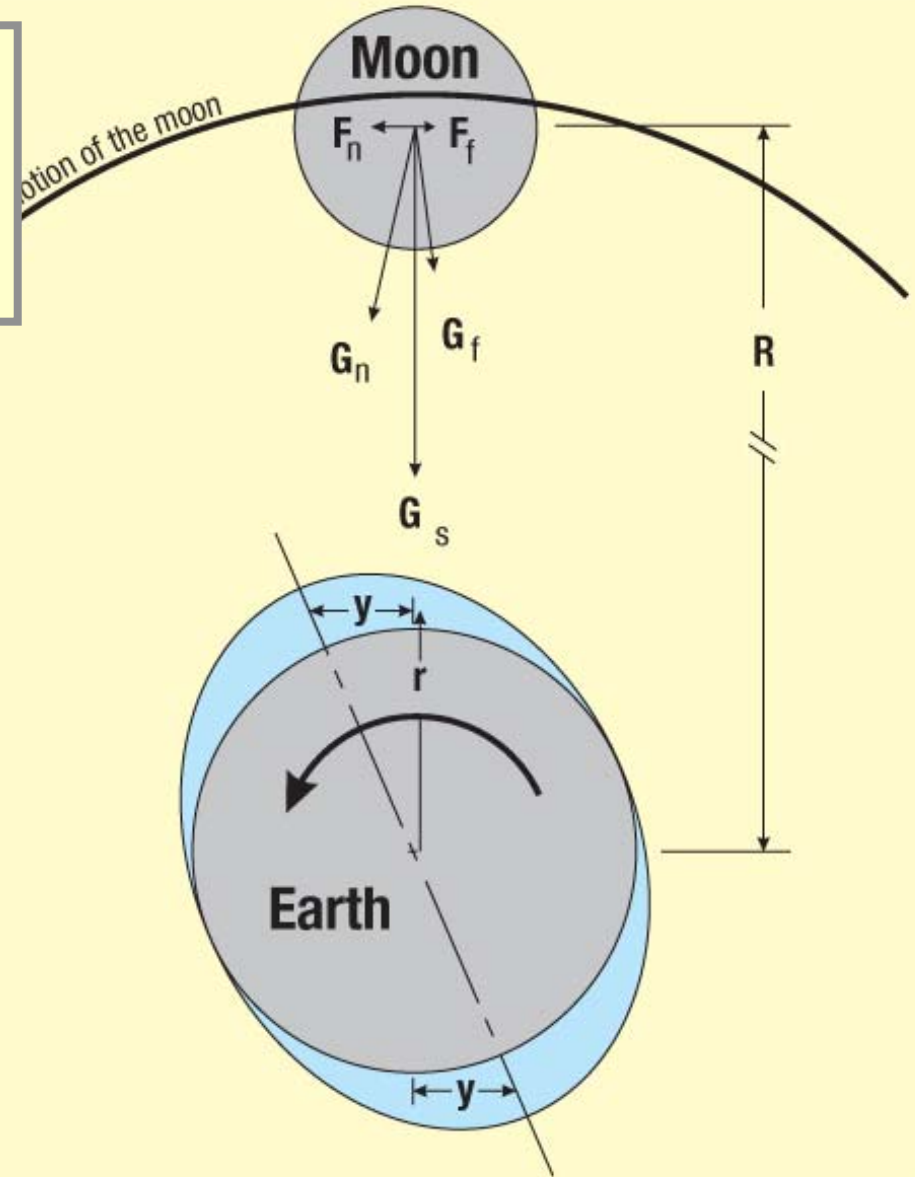
# Revolution without rotation



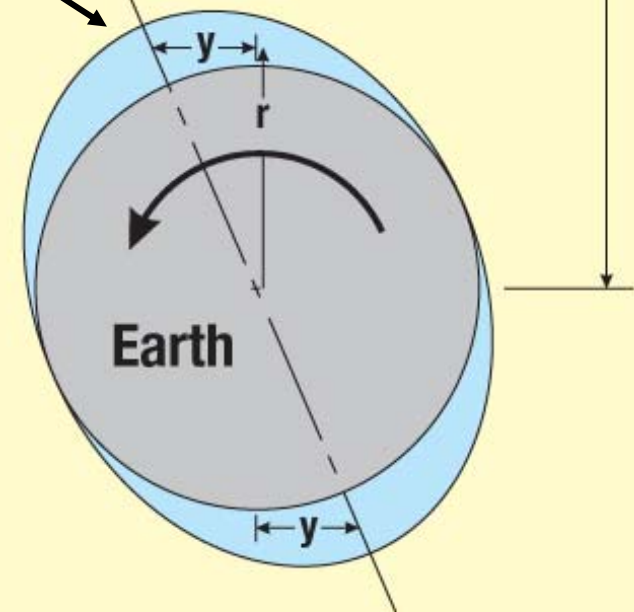
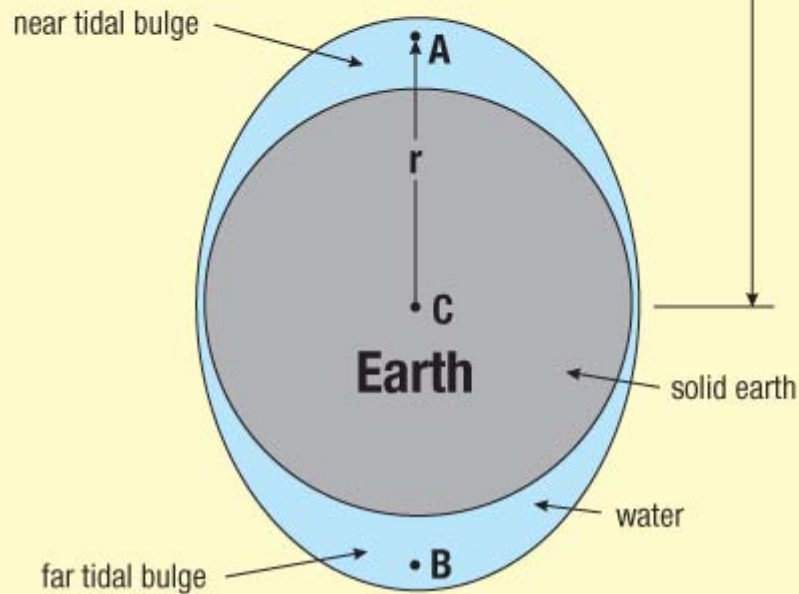
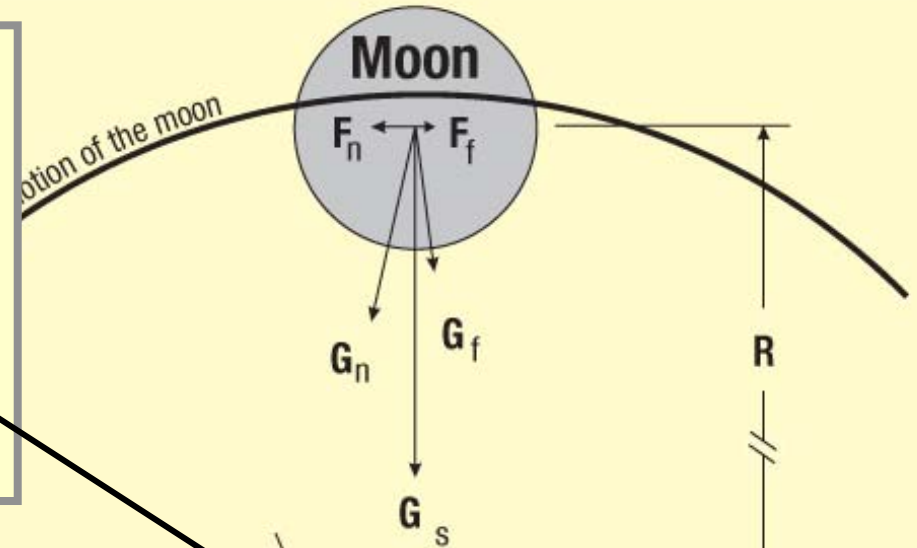
# But the Earth rotates!



- The Earth rotates faster than the Moon revolves

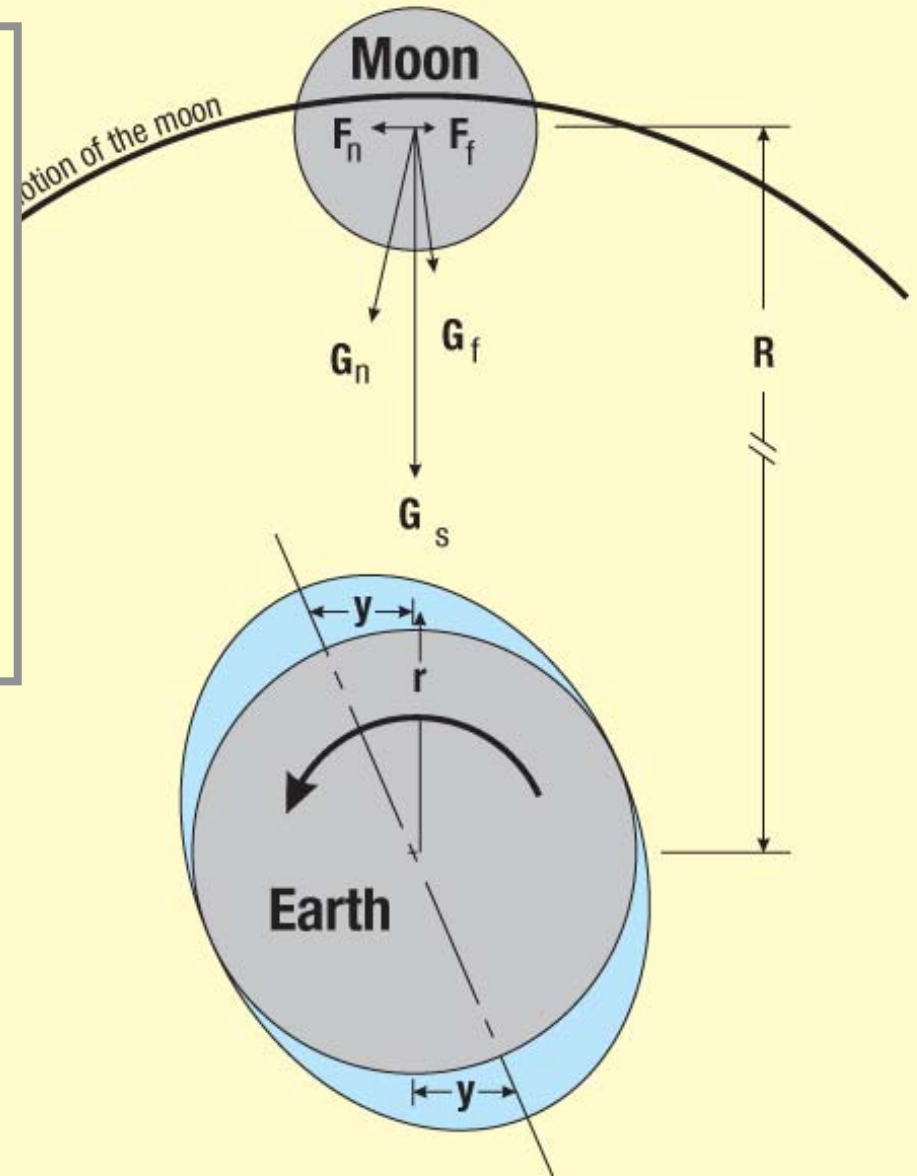
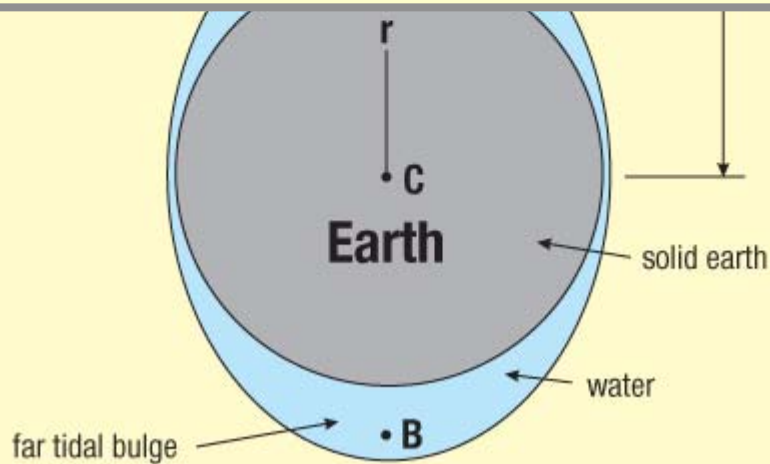


- The Earth rotates faster than the Moon revolves
- The tidal bulge is advancing the direct LOS

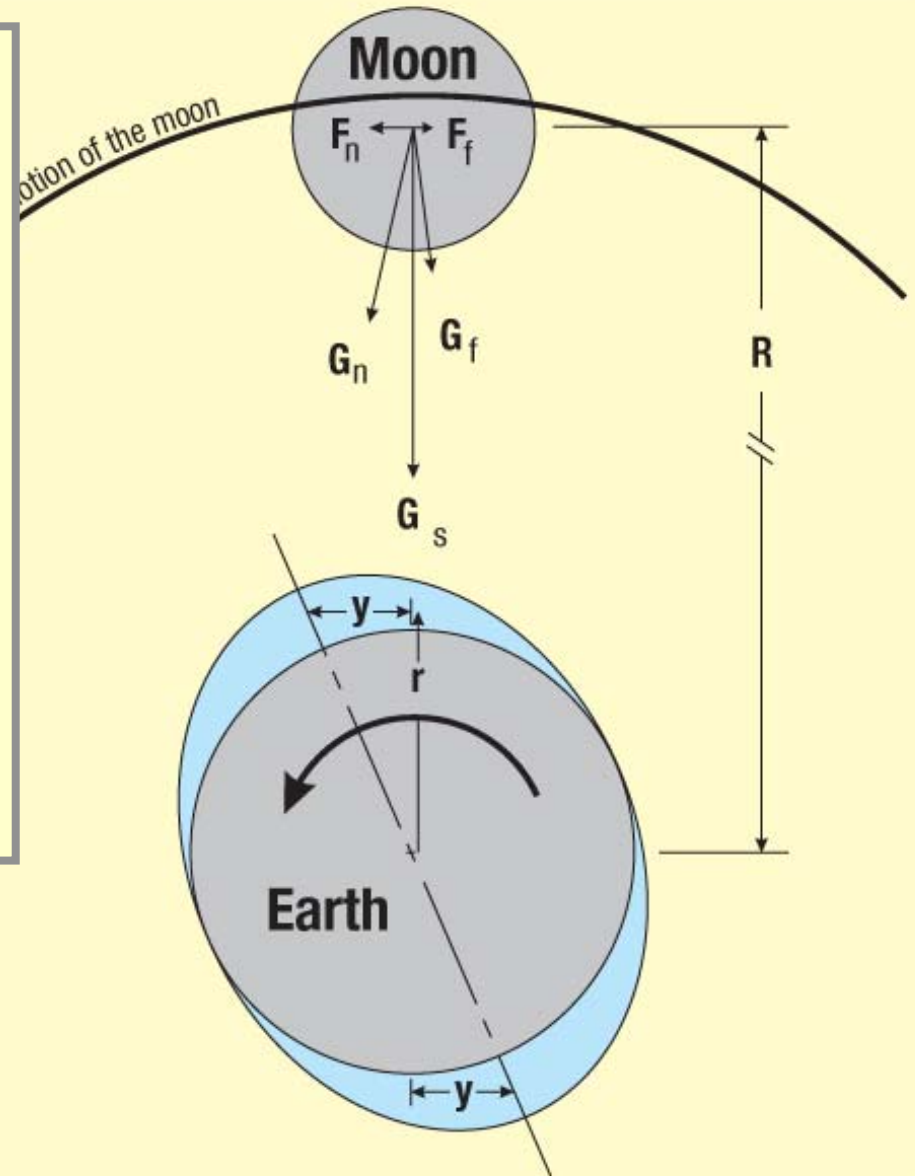
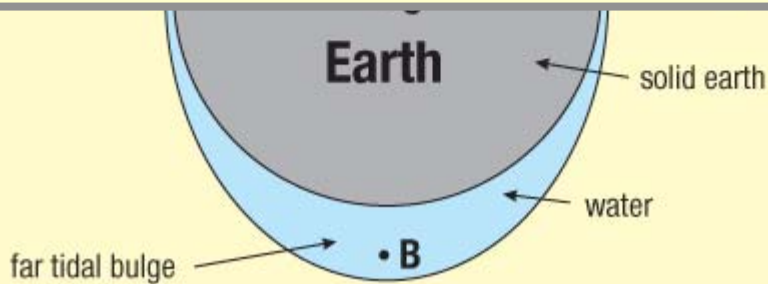




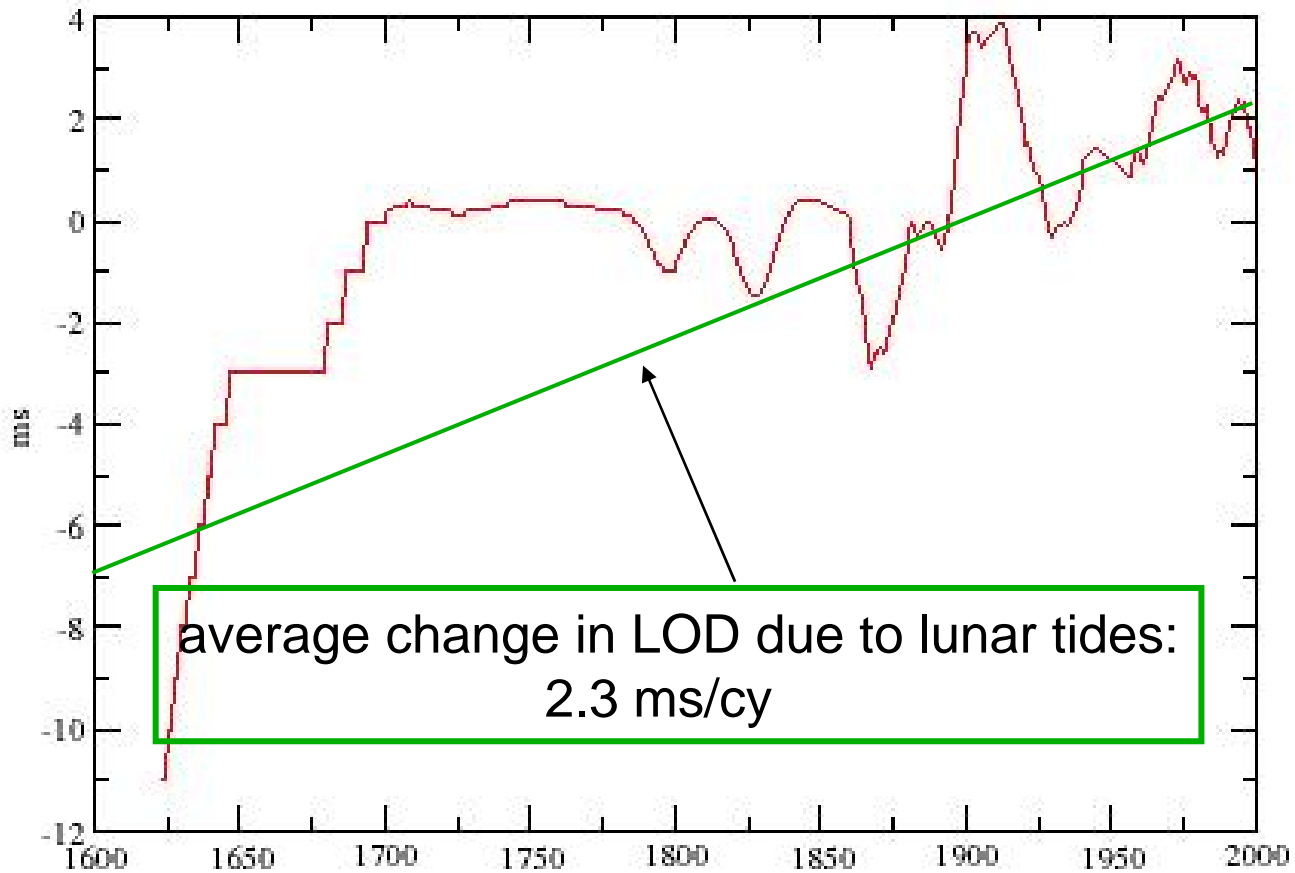
- The Earth rotates faster than the Moon revolves
- The tidal bulge is advancing the direct LOS
- Torques try to pull the tidal bulge back into LOS



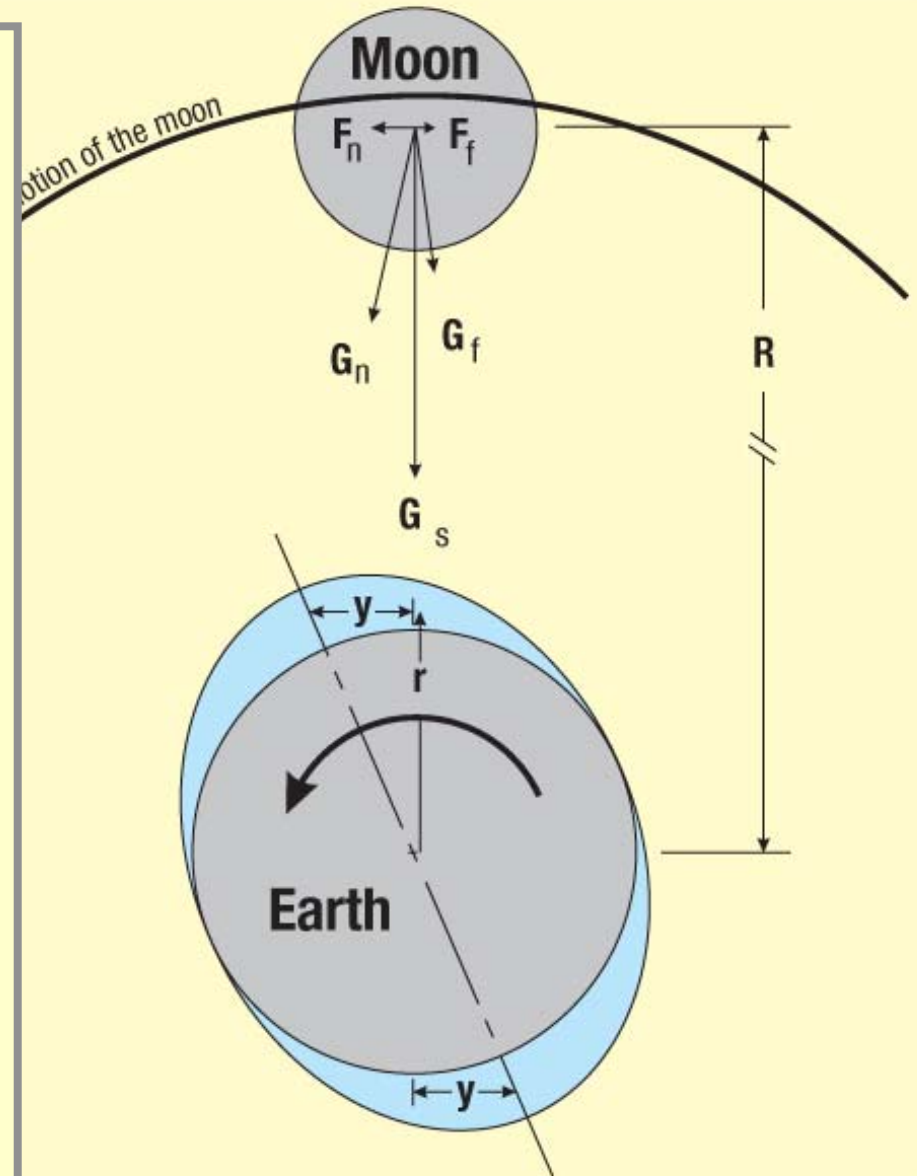
- The Earth rotates faster than the Moon revolves
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- Torques try to pull the tidal bulge back into LOS
- Consequence is a slow down of Earth rotation (tidal friction)



change in LOD relative to 86400 s



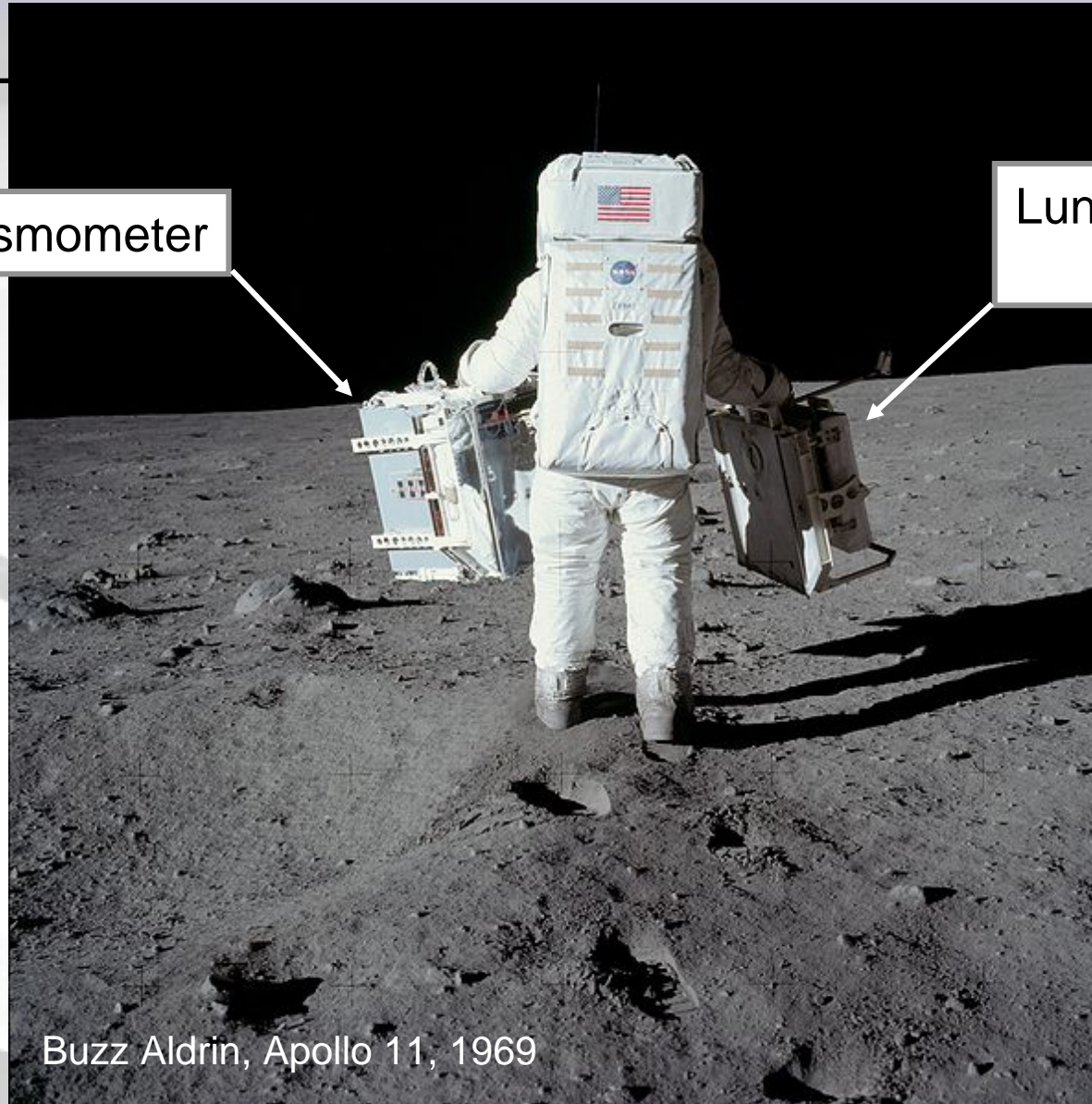
- The Earth rotates faster than the Moon revolves
- The tidal bulge is advancing the direct LOS
- Torques try to pull the tidal bulge back into LOS
- Consequence is a slow down of Earth rotation (tidal friction)
- another consequence from conservation of angular momentum: the Moon recedes from the Earth



# The first extraterrestrial applied space geophysicist

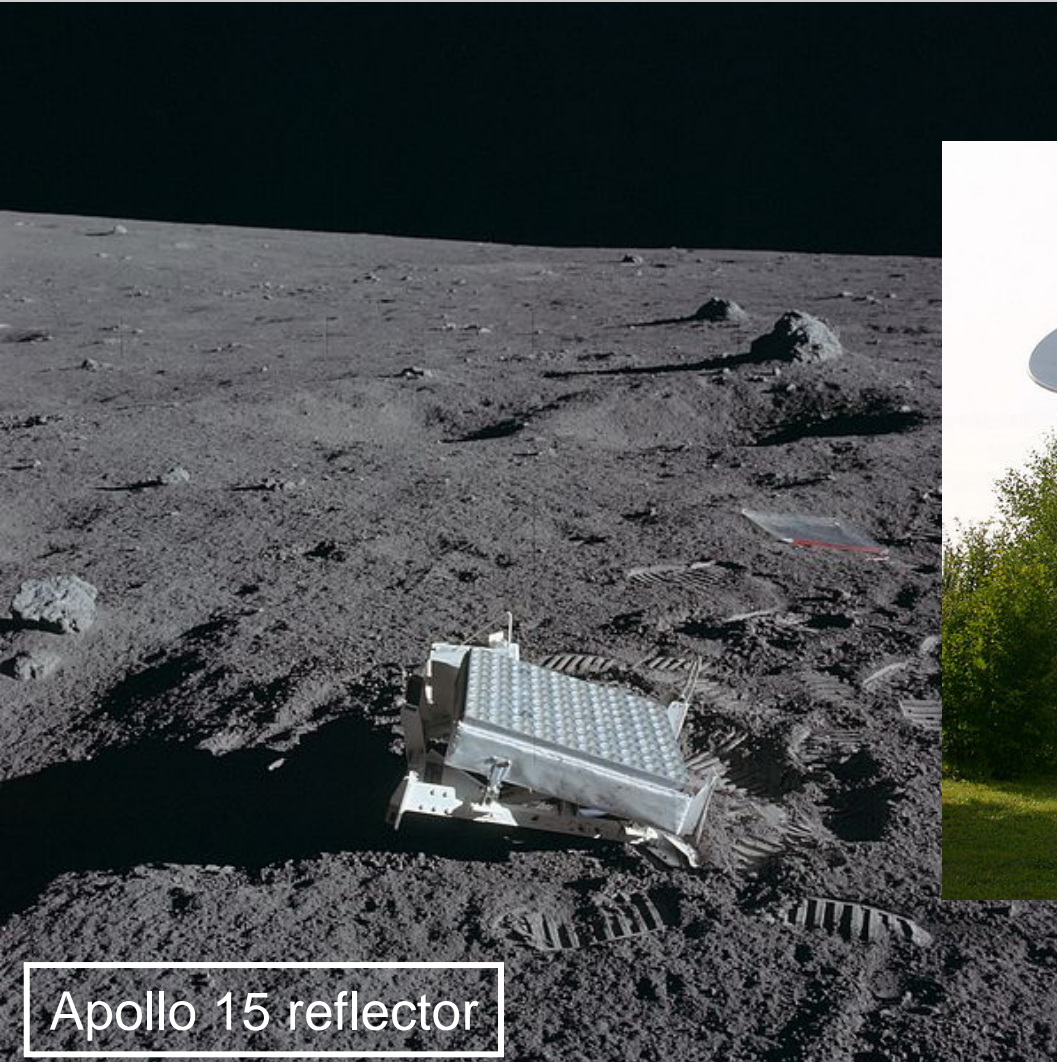
seismometer

Lunar Laser Ranging Reflector



Buzz Aldrin, Apollo 11, 1969

# Lunar Laser Ranging



Apollo 15 reflector



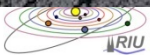
Wettzell Lunar Laser Station

# Lunar Laser Ranging

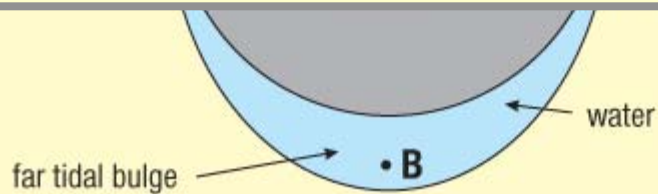
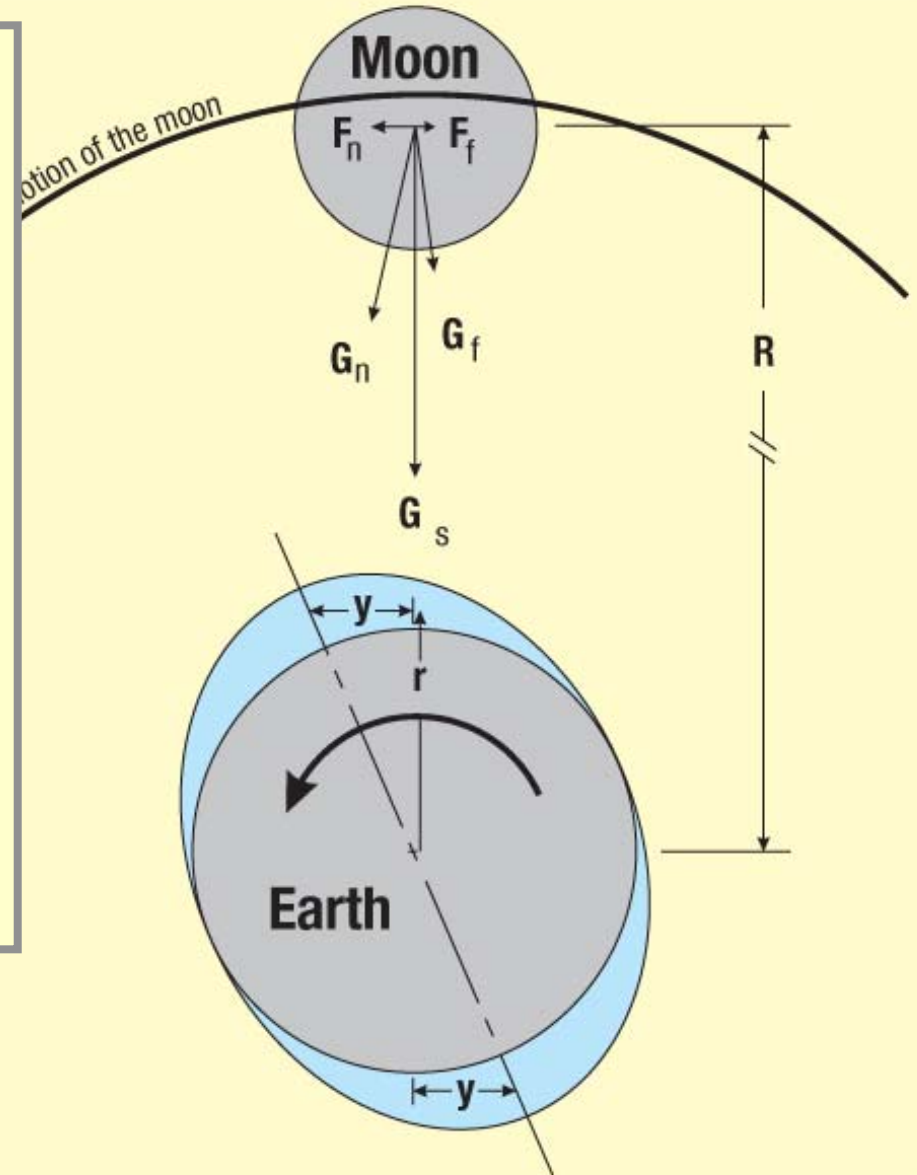
The Moon recedes from Earth by 4m/cy

Apollo 15 reflector

Wetzell Lunar Laser Station

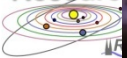


- The Earth rotates faster than the Moon revolves
- The **tidal bulge is advancing the direct LOS by  $3^\circ$**
- Earth rotates below the tidal bulge
- A location on the surface experiences periodically high and low tides





Ocean tides are the directly visible manifestation of tidal forces



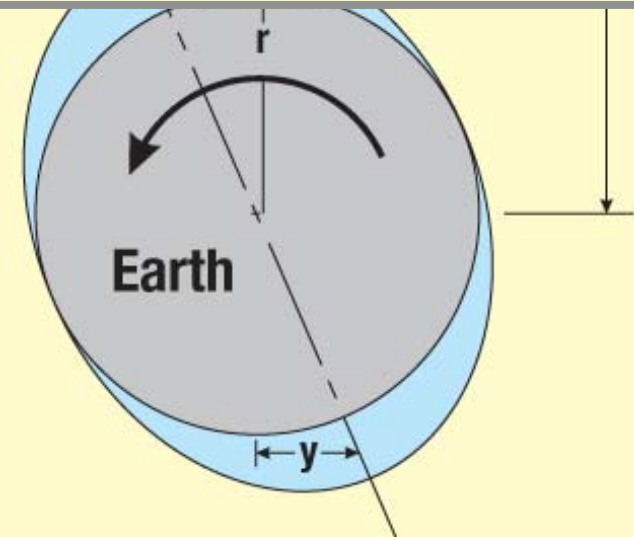
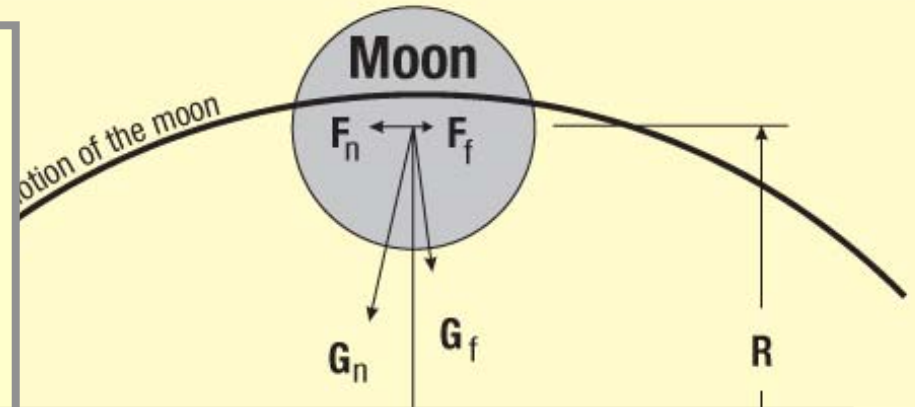
- **The Earth rotates faster than the Moon revolves**

- The tidal bulge is advancing the direct LOS

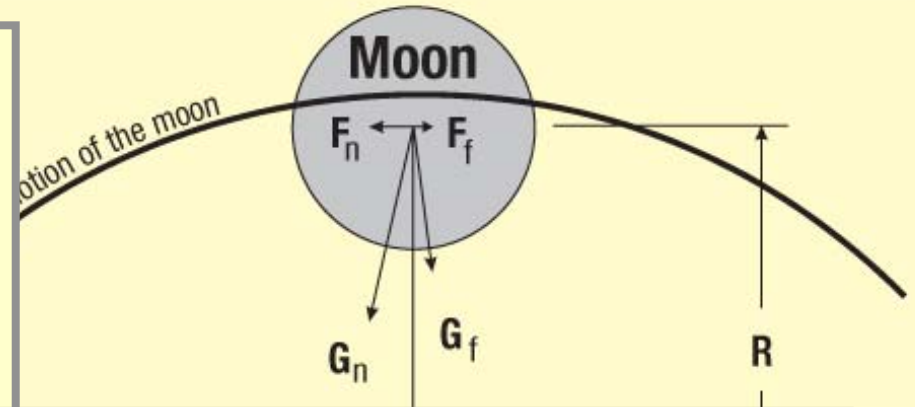
- **Is there a case in the solar system where a moon revolves faster than the planet rotates?**

- consequence of conservation of angular momentum of Earth rotation (tidal friction)

- another consequence from conservation of angular momentum: the Moon recedes from the Earth



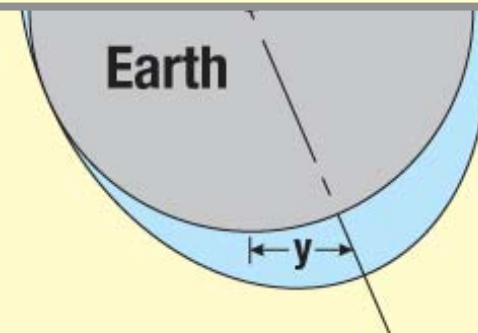
- **The Earth rotates faster than the Moon revolves**
- The tidal bulge is advancing the direct LOS



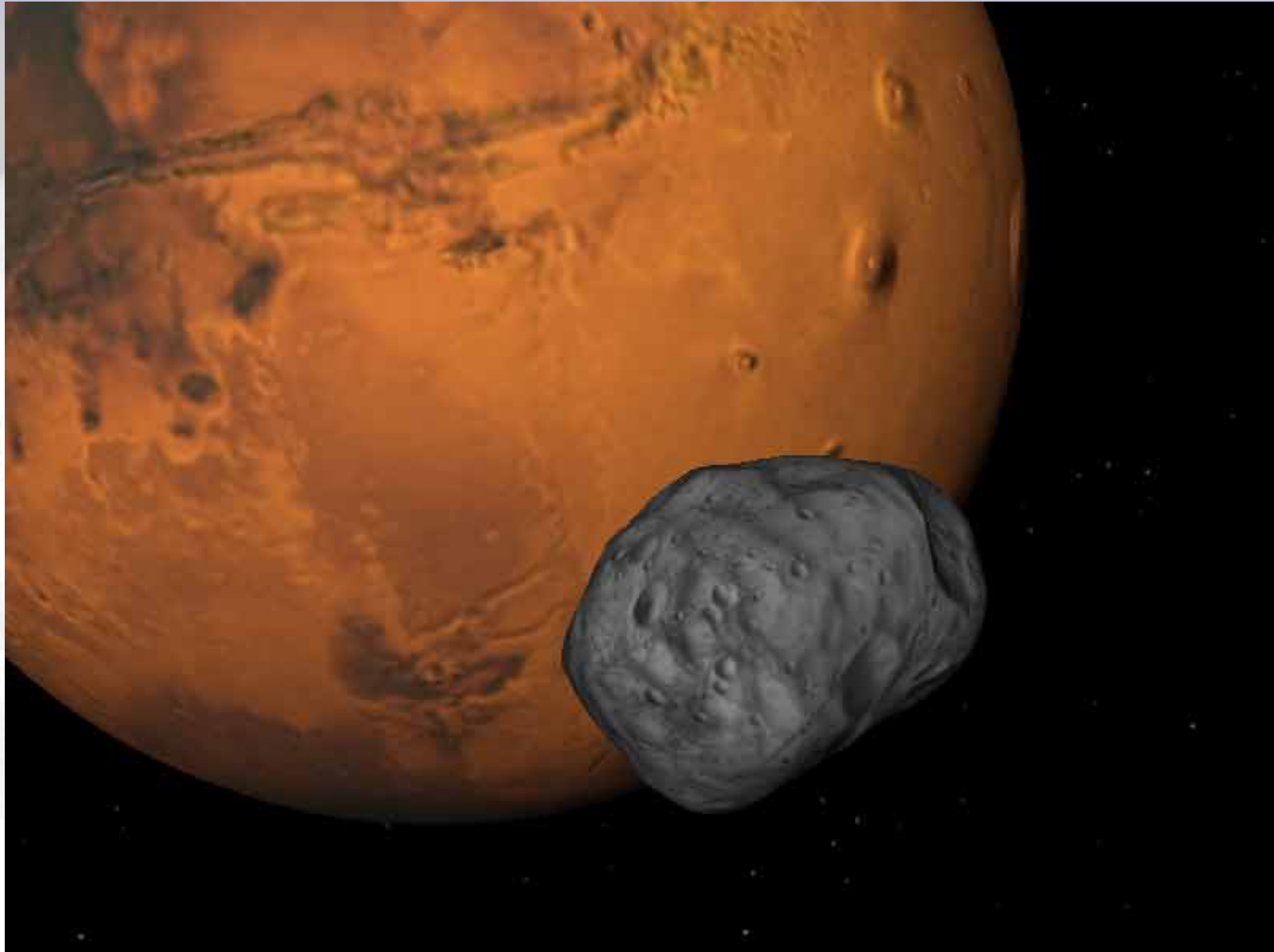
Is there a case in the solar system where a moon revolves faster than the planet rotates?

**YES, there is!**

- another consequence from conservation of angular momentum: the Moon recedes from the Earth



# The Mars/Phobos case



Rosetta\_CD\PR\what\_is\_RS.ppt, 10.06.2011 12:36AM, 28

# The Mars/Phobos case

## Phobos

size : 26 x 22 x 18 km  
period : 7 h 39 m  
distance : 9,378 km

## Mars

radius : 3,990 km  
period : 24 h 37 m

## Deimos

size : 15 x 12 x 10 km  
period : 30 h 17 m  
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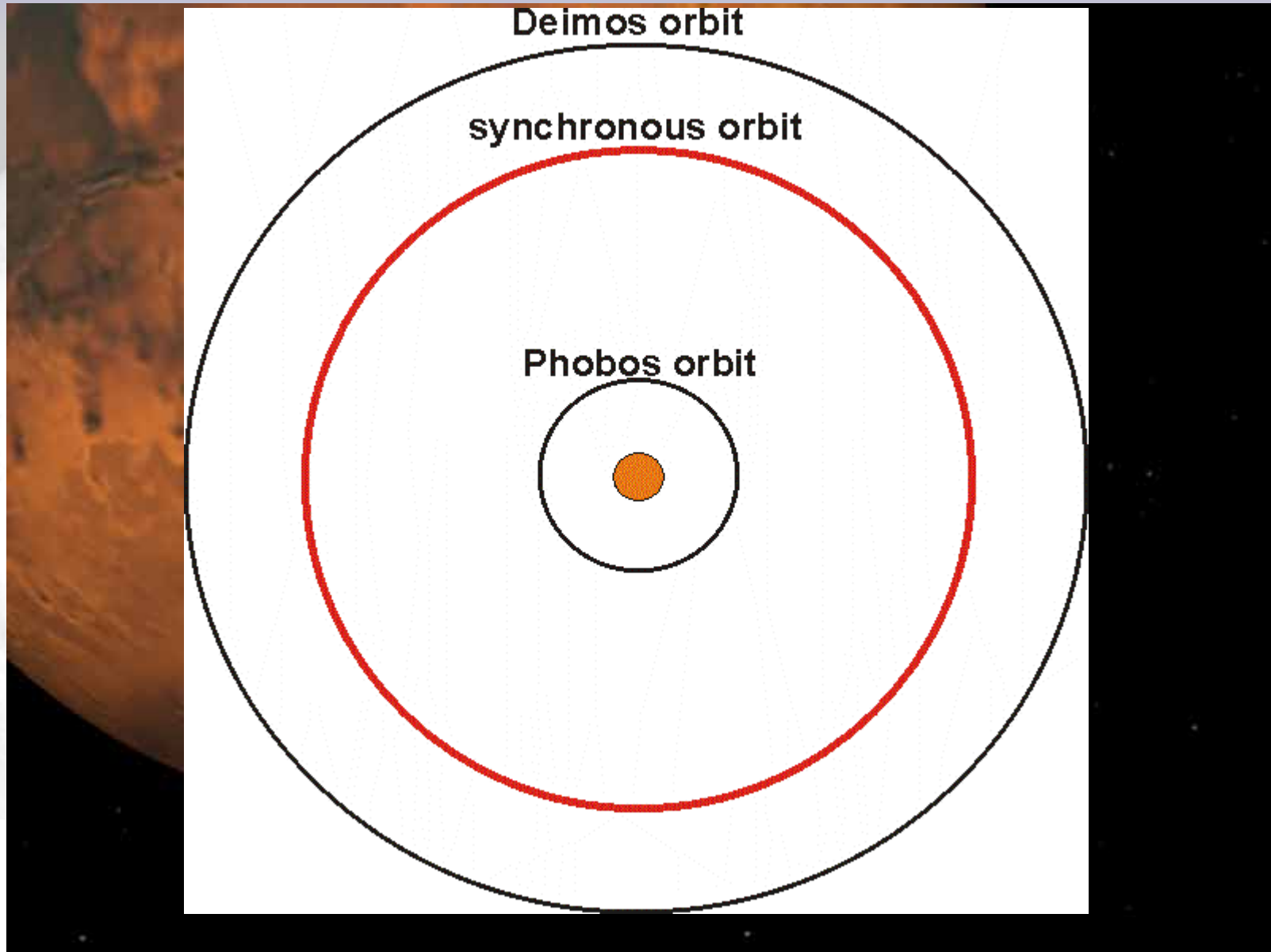
## Deimos

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synchronous orbit:  
planetary rotation = moon revolution

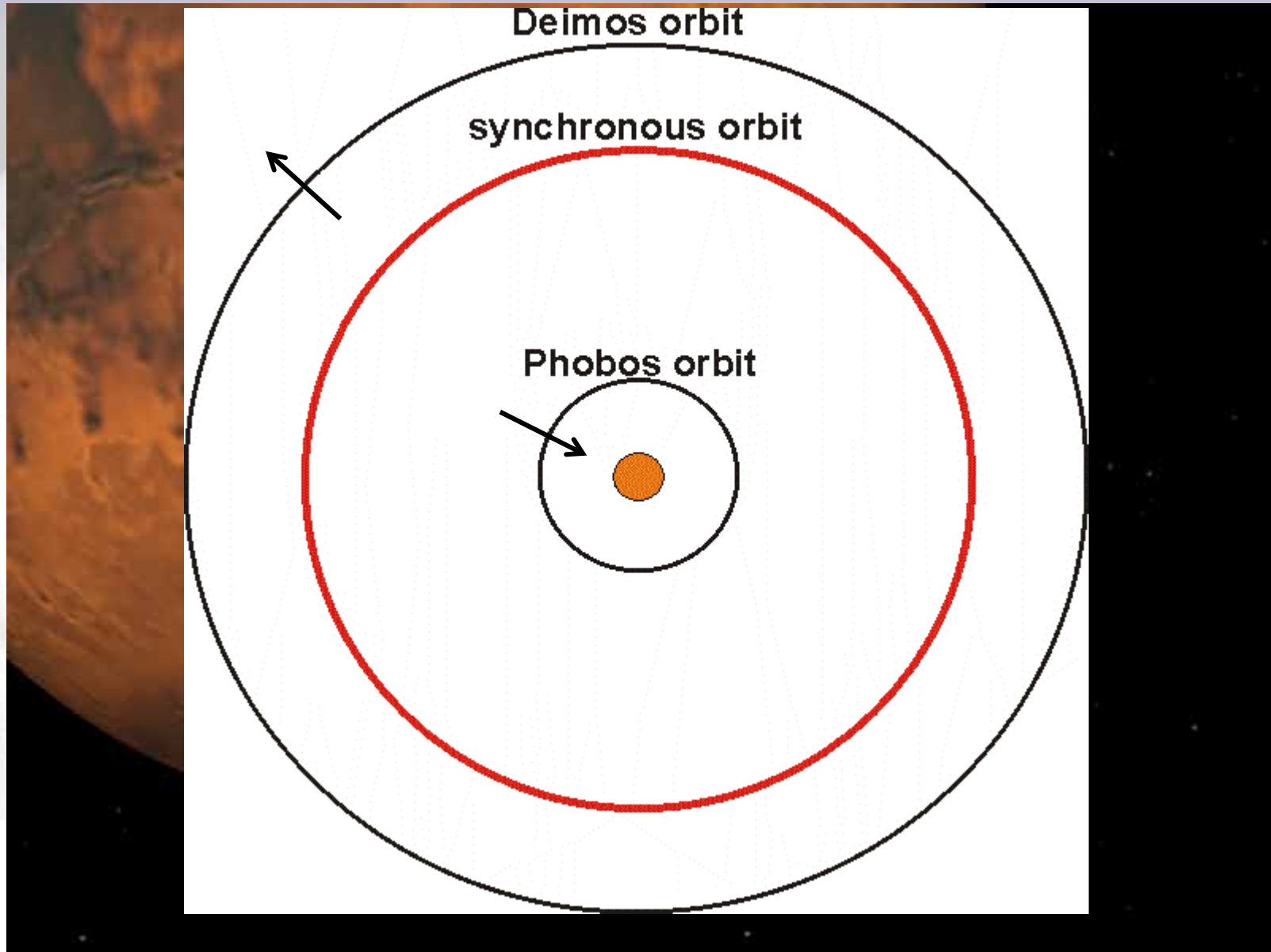
Mars  
24 h 37 m  
21,000 km distance  
Phobos: inside  
Deimos: outside

# The Mars/Phobos case



Rosetta\_CD\PR\what\_is\_RS.ppt, 10.06.2011 12:36AM, 31

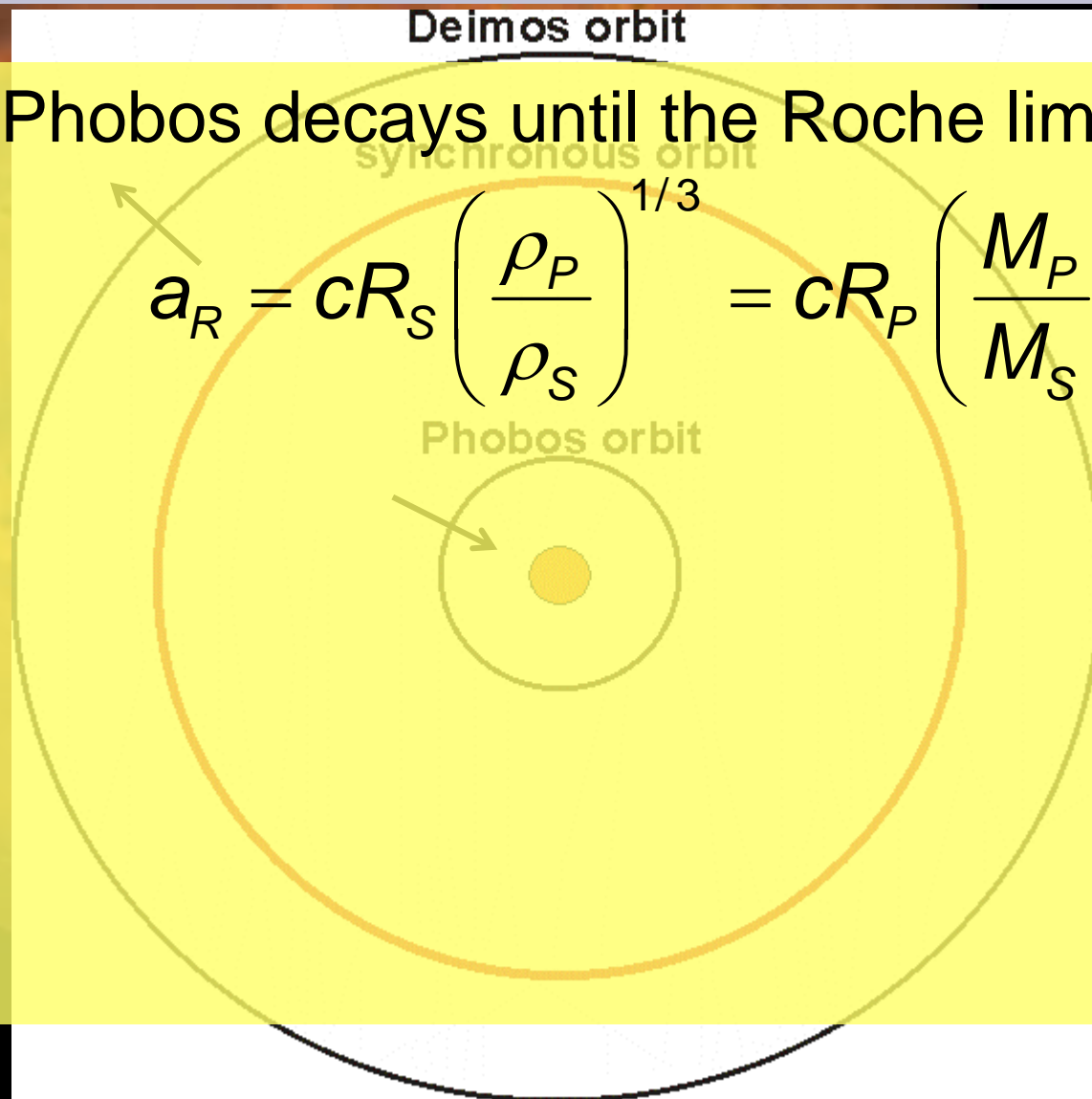
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Roche limit:

$$a_R = cR_S \left( \frac{\rho_P}{\rho_S} \right)^{1/3} = cR_P \left( \frac{M_P}{M_S} \right)^{1/3}$$

Proportionality factor  $c$  depends on the nature of the

secondary: fluid	$c = 2.44 \Rightarrow$	$a_R = 10,500$ km
solid	$c = 1.26 \Rightarrow$	$a_R = 5,500$ km
rubble pile	$c = 1.54 \Rightarrow$	$a_R = 6,700$ km

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Phobos is highly porous  $\Rightarrow$  „rubble pile“ (Andert et al., 2010)

When will Phobos reach the Roche limit?

$$\tau_{\text{Roche}} = \frac{\frac{2}{13} \left( a^{13/2} - a_R^{13/2} \right)}{3 \frac{k_{2,\text{Mars}}}{Q_{\text{Mars}}} \frac{M_{\text{Ph}}}{M_{\text{Mars}}} R_{\text{Mars}}^5 \sqrt{GM_{\text{Mars}}}}$$

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## Other tidal phenomena in the solar system

---

- forcing Mercury in a 3/2 resonance
- forcing all moons in the solar system into synchronous rotation
- deforming the Jupiter moon Io periodically => tidal heating => volcanic activity
- and the non-plus-ultra:

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- deforming the Jupiter moon Io periodically => tidal heating => volcanic activity
- and the non-plus-ultra: synchronizing the Pluto system



# Tides in extrasolar systems

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- studying tides in extrasolar planetary systems requires the knowledge of mass and size of the planet and the star
- feasible with transiting planets combined with radial velocity observations and spectroscopy
- orbital decay => survivability
- stellar spin-up
- spin-orbit coupling => synchronization of rotation and revolution
- circularization =>  $e \rightarrow 0$

# The tidal potential or the Doodson constant

two tidal potentials:

- a) The primary acting on the secondary
- b) The secondary acting on the primary

$$D_{ps} = \frac{GM_p}{a^3} R_s^2$$

$$D_{sp} = \frac{GM_s}{a^3} R_p^2$$

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Earth/Moon system:

Earth on Moon  
 $D_{ps} = 20.9 \text{ m}^2/\text{s}^2$

Moon on Earth  
 $D_{sp} = 3.5 \text{ m}^2/\text{s}^2$

Sun on Earth  
 $D_{sp} = 1.6 \text{ m}^2/\text{s}^2$

When will the planet reach the stellar Roche limit?

$$\tau_{\text{Roche}} = \frac{\frac{2}{13} \left( a^{13/2} - a_R^{13/2} \right)}{3 \frac{k_{2,star}}{Q_{star}} \frac{M_P}{M_{star}} R_{star}^5 \sqrt{GM_{star}}}$$

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stellar property factor

When will the planet reach the stellar Roche limit?

$$\tau_{\text{Roche}} = \frac{\frac{2}{13} \left( a^{13/2} - a_R^{13/2} \right)}{3 \frac{k_{2,star}}{Q_{star}} \frac{M_P}{M_{star}} R_{star}^5 \sqrt{GM_{star}}}$$

- 
- tidal forces affect the evolution of extrasolar planetary systems
    - if  $a < 0.1$  AU
    - if  $M_p$  is large
    - may force orbits towards  $e \rightarrow 0$
    - may lead to the loss of companions
    - may force massive companions into double-synchronous rotation
      - but the companion is not safe....
    - may spin-up stellar rotation (look for old fast rotating stars)

...book your next boat tour  
when the tide is in!

